

Bars and galaxy interactions in the Illustris simulation

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with E. Lokas

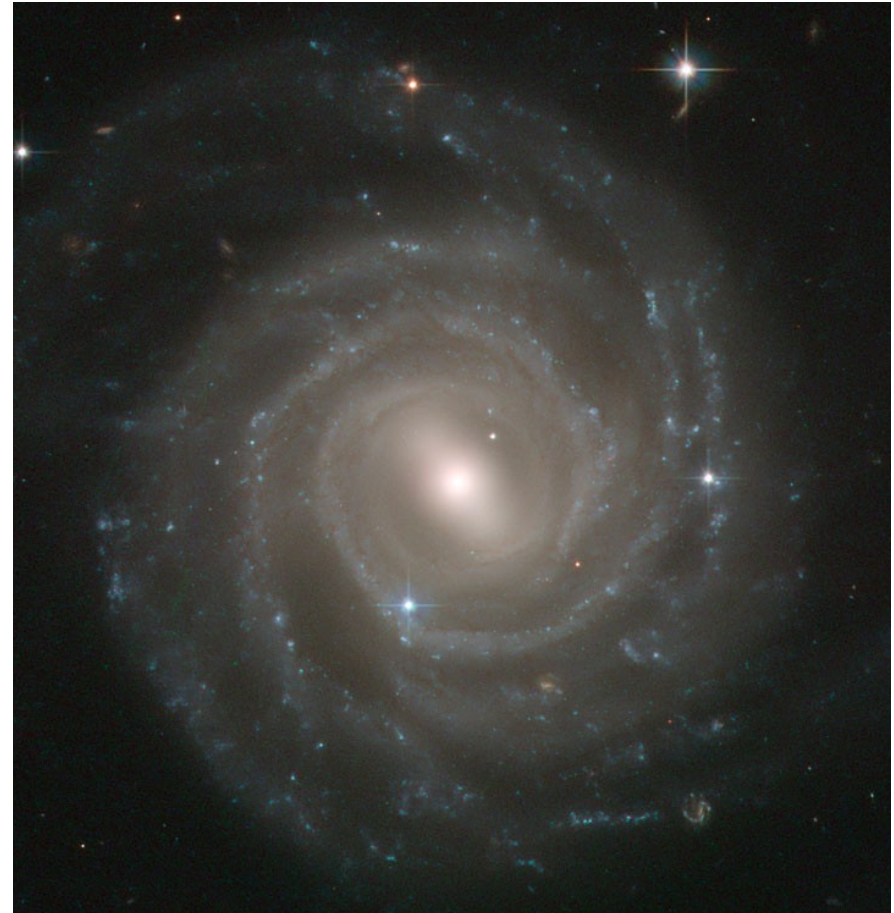
Nicolaus Copernicus Astronomical Center, Warsaw

PNCG, Lyon, November 17, 2017



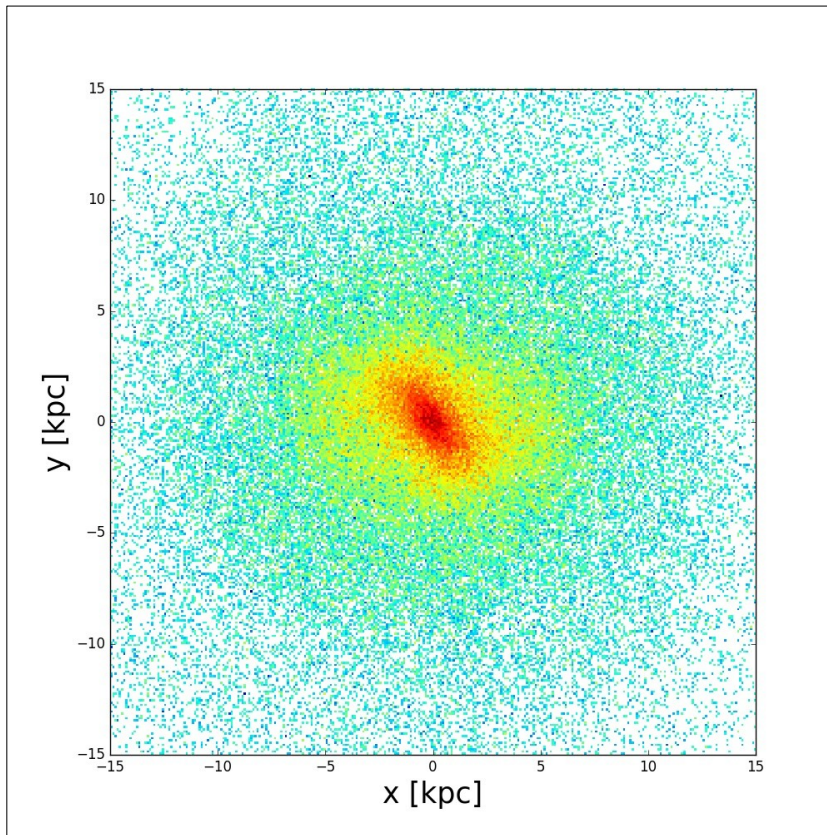
Introduction

- 30-60% of disc galaxies in the local universe host a bar in their center (Barazza et al. 2008, Nair & Abraham 2010, Laine et al. 2016)
- Bars can be formed in situ, or triggered by interactions with other galaxies: Tidal Bars
- Illustris : high resolution hydrodynamical cosmological simulation : $(106 \text{ Mpc})^3$, ~ 18 billion particles (Volgelsberger et al. 2015)

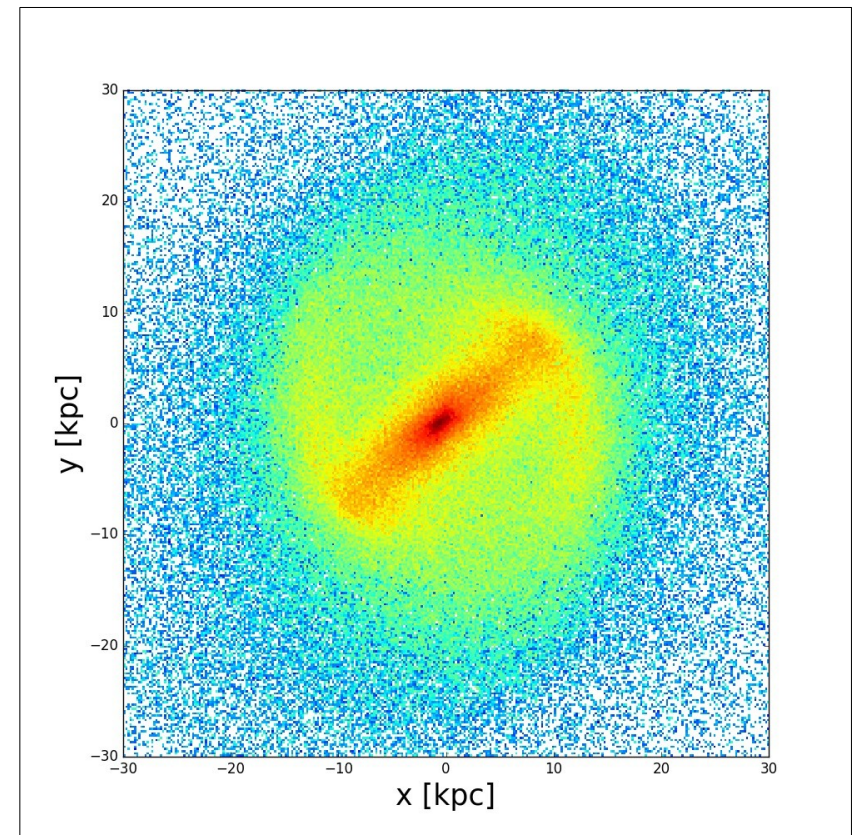


UGC 12158, Hubble Space Telescope

Bars in Illustris galaxies



Weak Bar



Strong Bar

Bars in Illustris galaxies

- Definition of bar strength :

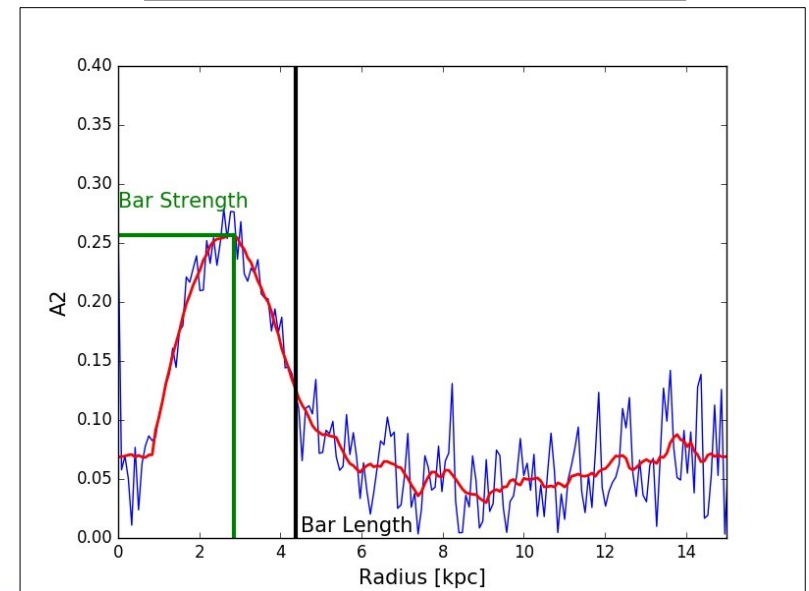
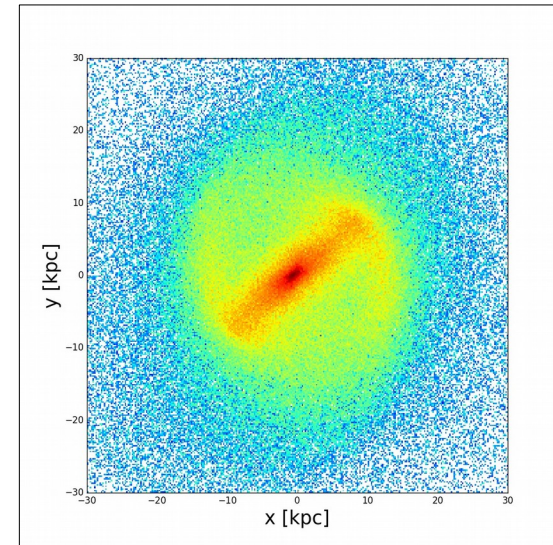
Fourier moments:

$$- a_m(R) = \sum_i M_i \cos(m \phi_i)$$

$$- b_m(R) = \sum_i M_i \sin(m \phi_i)$$

$$\text{Amplitude of } m=2 \text{ component } A_2: A_2(R) = \frac{\sqrt{a_2^2 + b_2^2}}{a_0}$$

→ Bar Strength: $A_{2,\max} = \max(A_2)$



Bars in Illustris galaxies

- Definition of bar strength :

Fourier moments:

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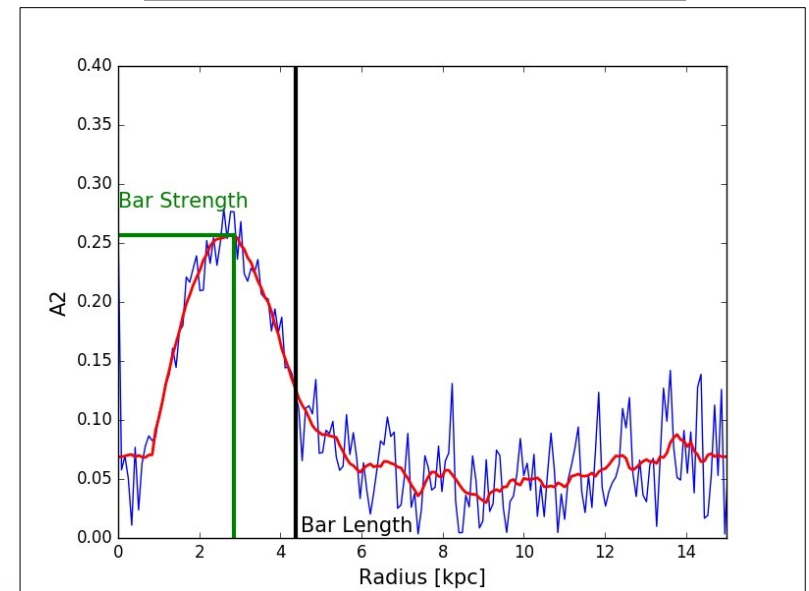
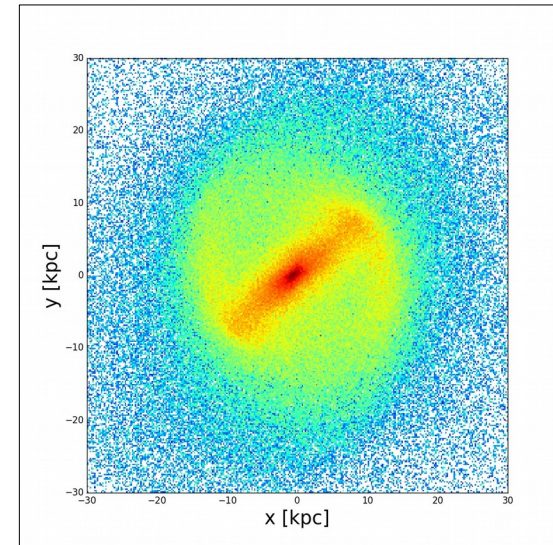
→ Bar Strength: $A_{2,\max} = \max(A_2)$

- Presence of a Bar if $A_{2,\max} > 0.15$:

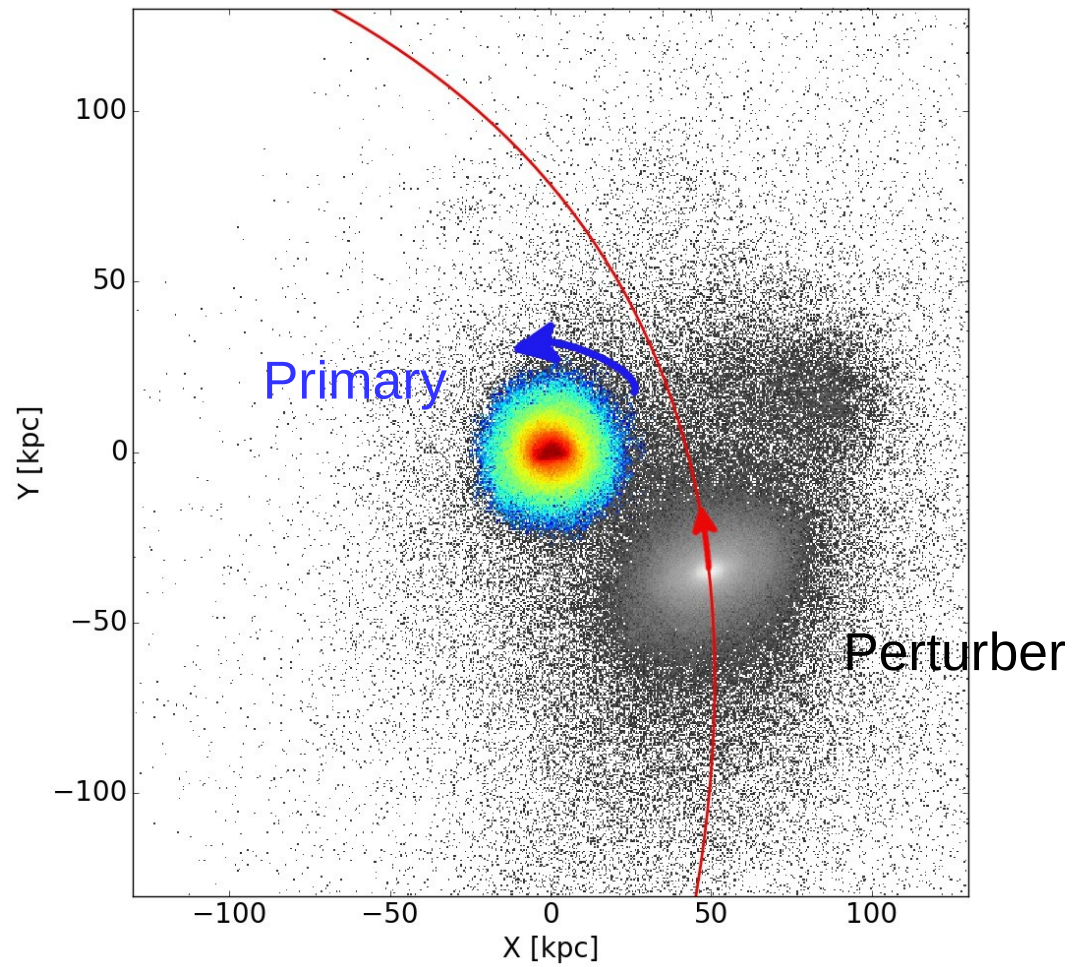
→ ~26% of bars in local disc galaxies

→ bar fraction ~ constant up to redshift 1

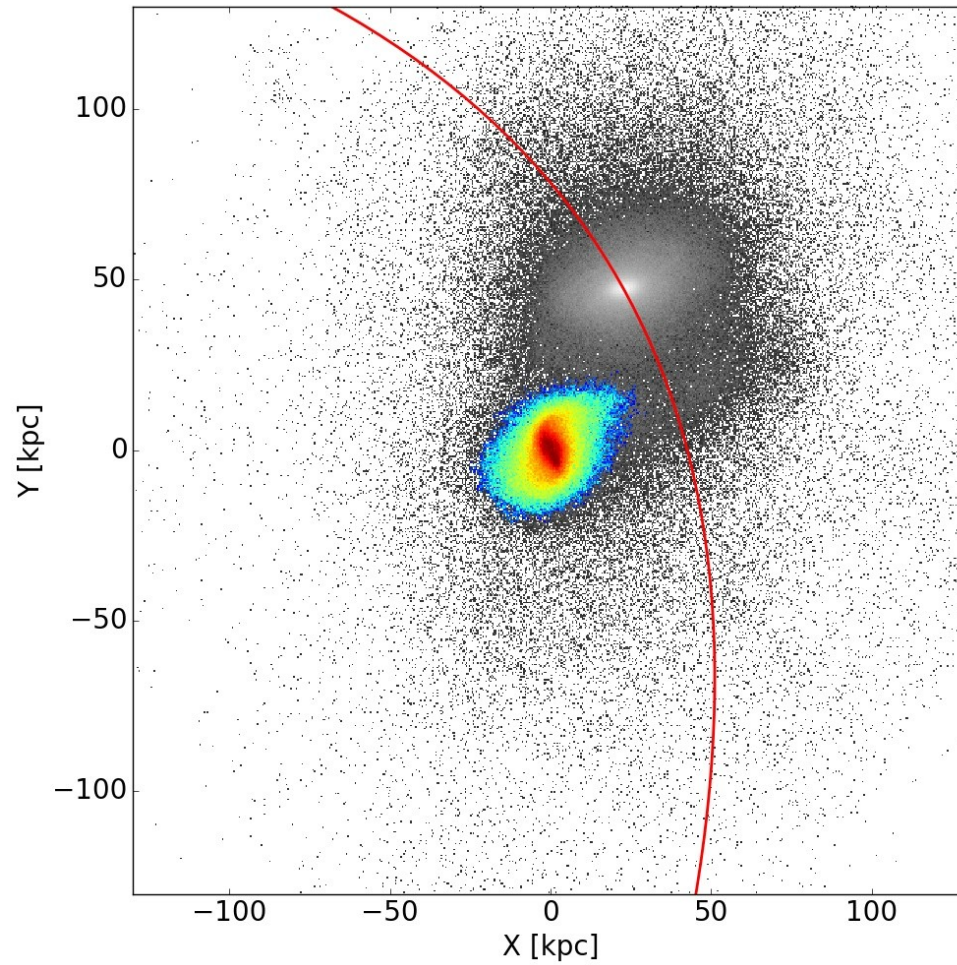
→ bar fraction increases with total mass



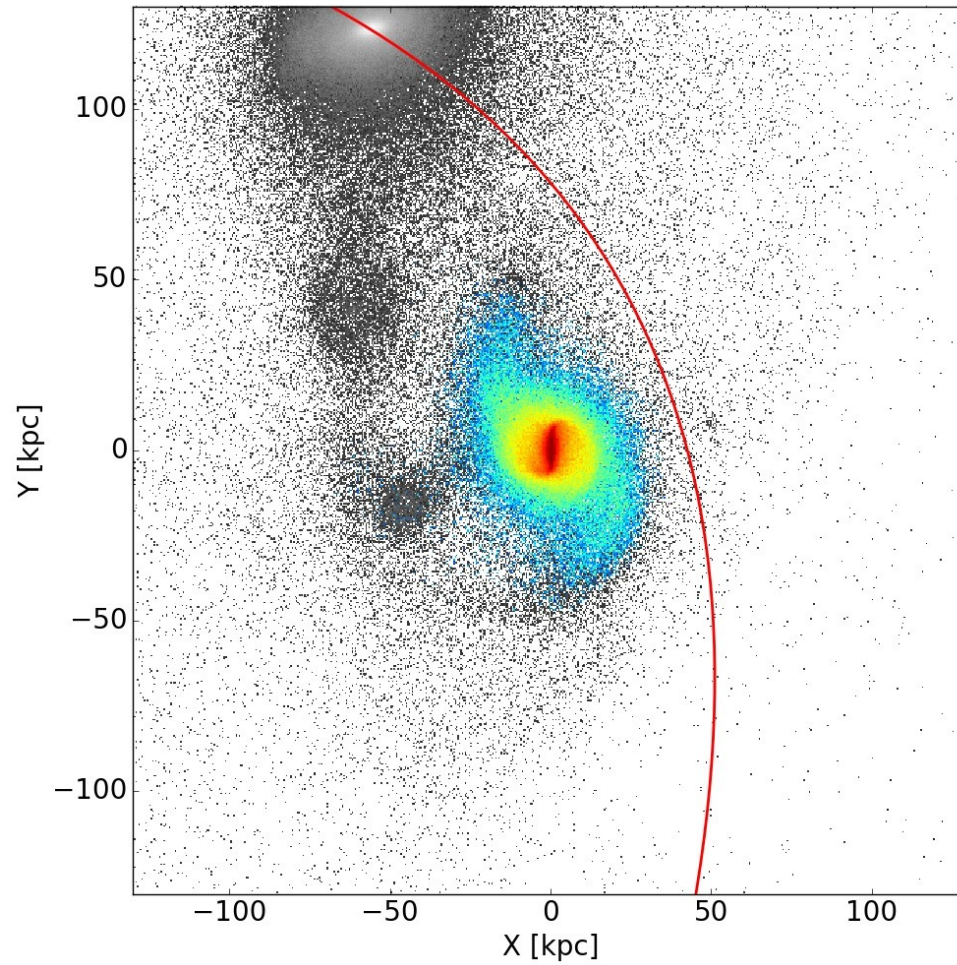
Bars and interactions



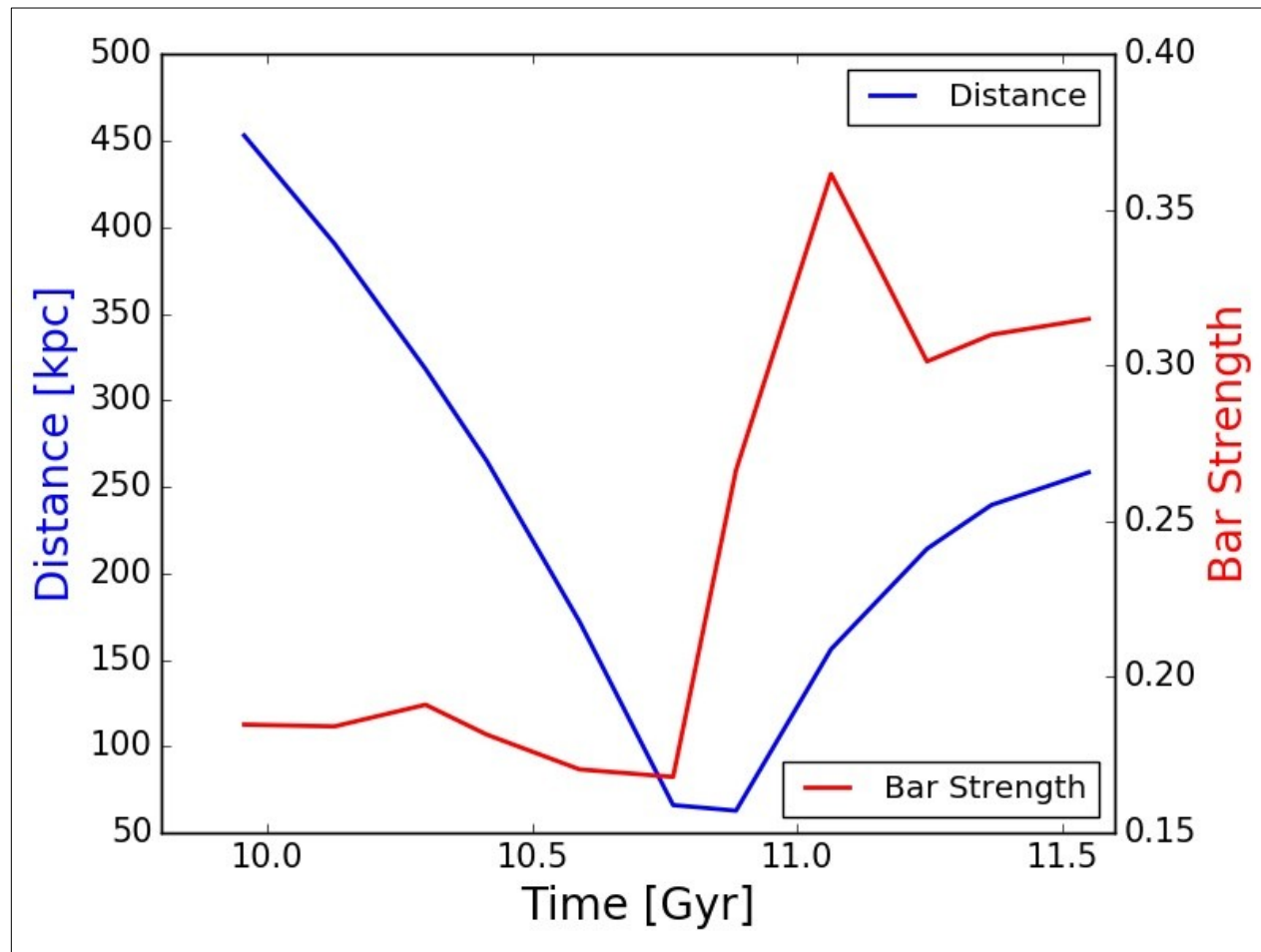
Bars and interactions



Bars and interactions

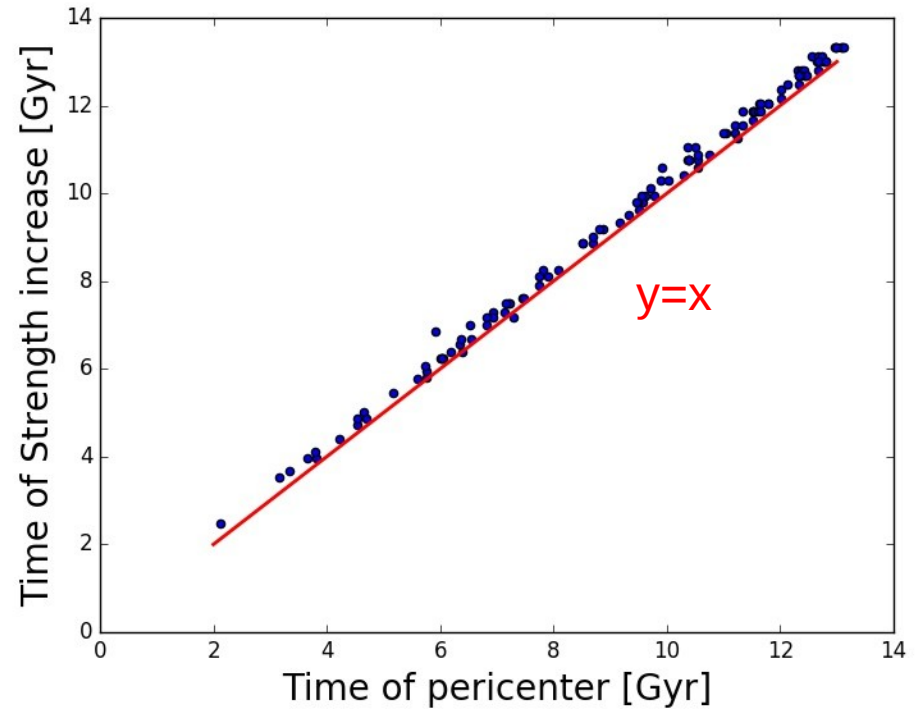


Bars and interactions



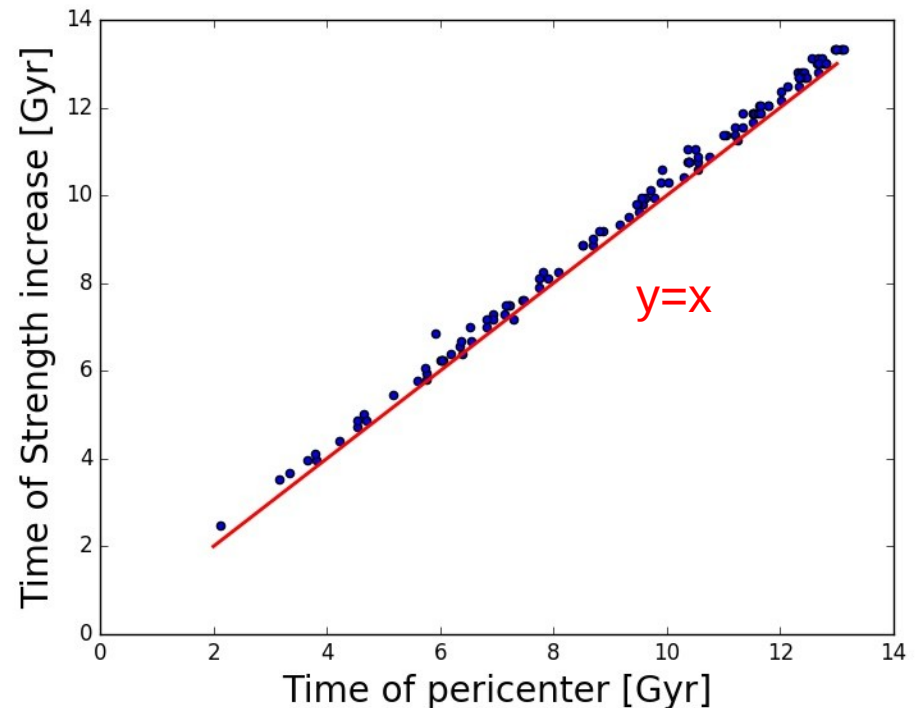
Bars and interactions

- Sample of 122 galaxies undergoing a fly-by interaction
- Bar strength increase happens right after the pericenter passage (0.29 ± 0.14 Gyr)



Bars and interactions

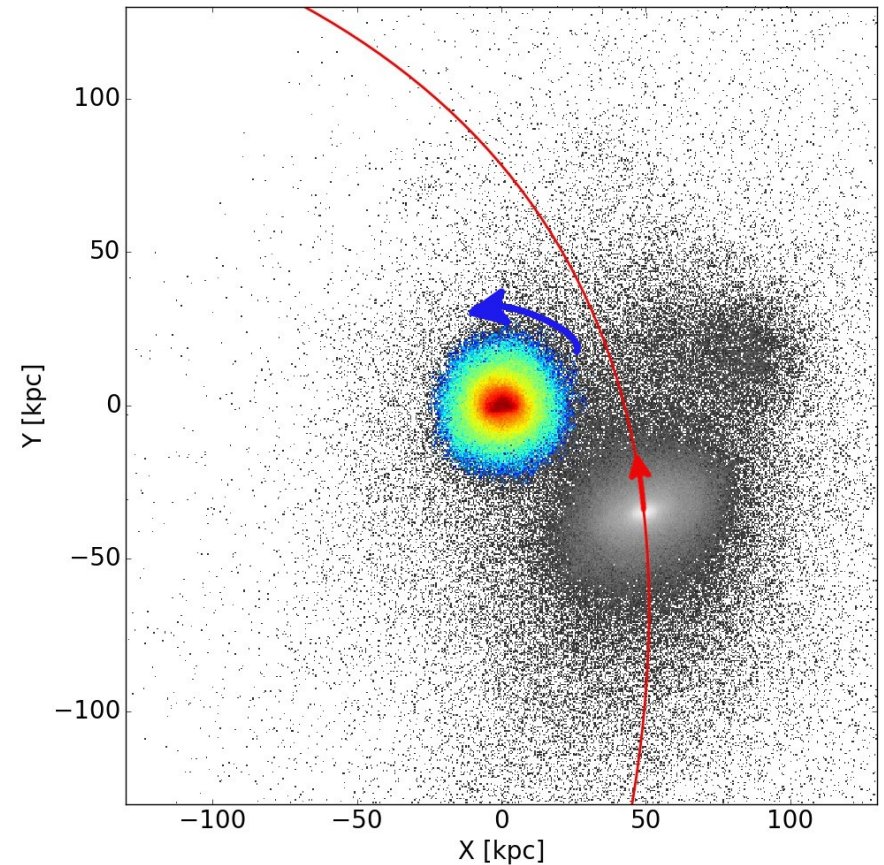
- Sample of 122 galaxies undergoing a fly-by interaction
- Bar strength increase happens right after the pericenter passage (0.29 ± 0.14 Gyr)
- In case of a pre-existing bar, the fly-by can:
 - Increase the bar strength
 - Decrease the bar strength
 - Have no effect
- Which parameters are responsible for these differences?



Bars and interactions

Orbital Angle

- Orbital Angle θ : Angle between the orbital plane of the encounter, and the primary disc plane:
 - Prograde encounter:
 $\cos(\theta) \sim 1$
 - Retrograde encounter:
 $\cos(\theta) \sim -1$

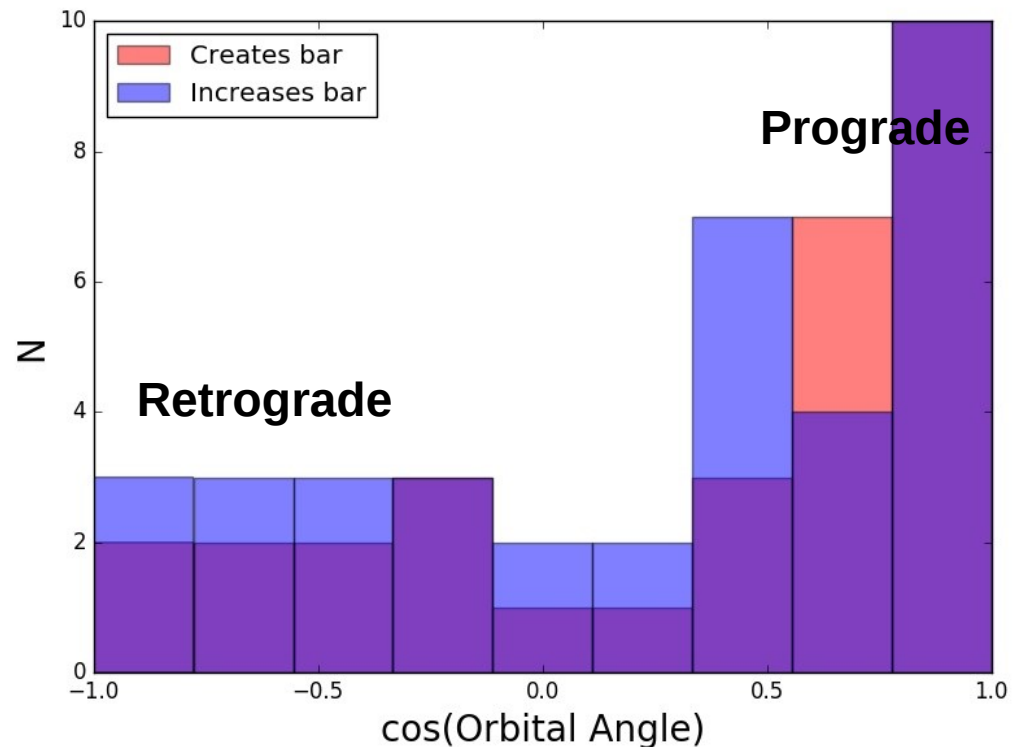


Prograde case

Bars and interactions

Orbital Angle

- Orbital Angle θ : Angle between the orbital plane of the encounter, and the primary disc plane:
 - Prograde encounter:
 $\cos(\theta) \sim 1$
 - Retrograde encounter:
 $\cos(\theta) \sim -1$
- Mostly pro-grade encounters create bars, or increase the bar strength



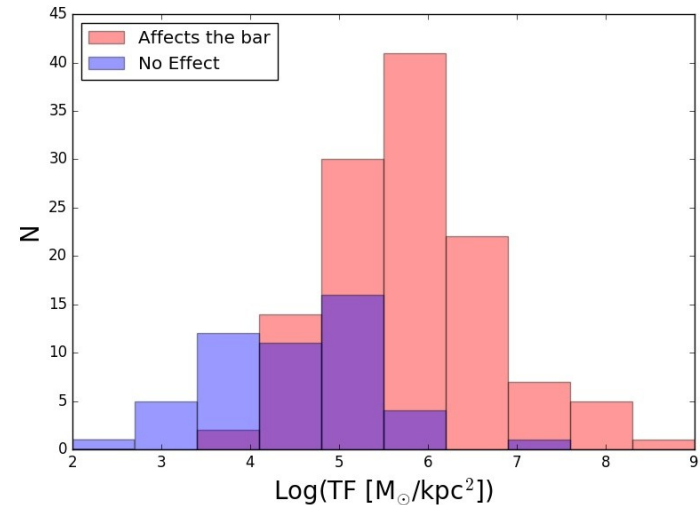
Bars and interactions

Tidal Force

- Strength of the interaction: Tidal Force (Lokas et al. 2016):

$$TF = \frac{R_1 * M_2}{D^3}$$

R1: scale-length of primary galaxy
M2: mass of the perturber
D: distance between the 2 galaxies



Bars and interactions

Tidal Force

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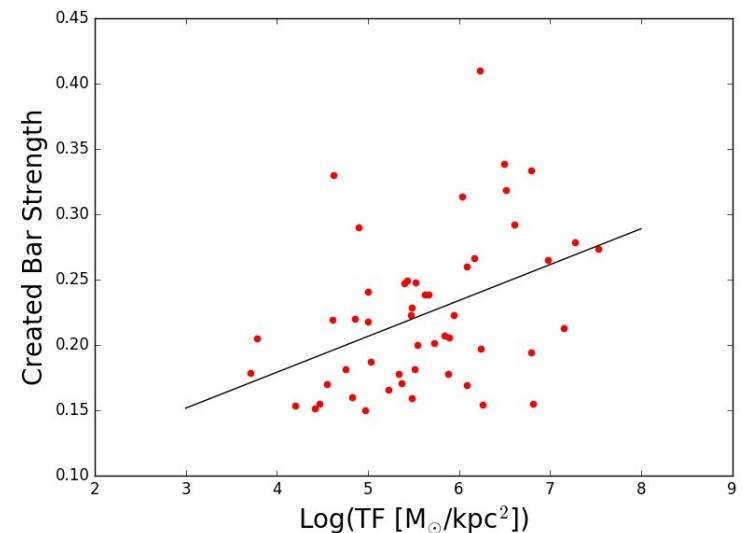
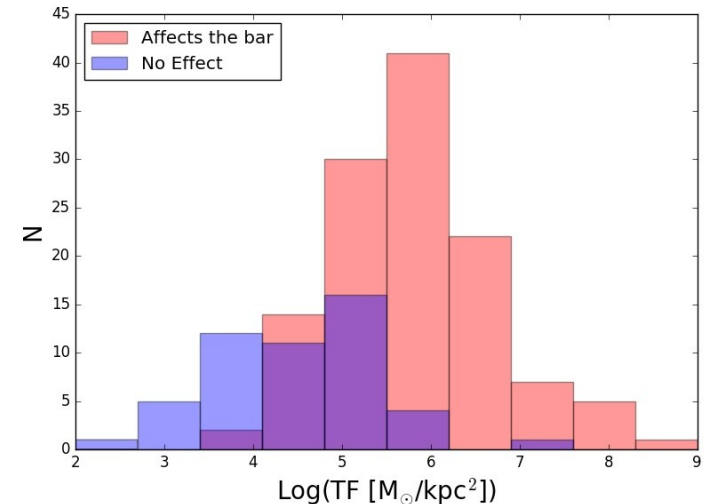
$$TF = \frac{R_1 * M_2}{D^3}$$

R1: scale-length of primary galaxy

M2: mass of the perturber

D: distance between the 2 galaxies

- The stronger the interaction, the stronger the created bar: *weak* correlation
- Many other parameters can come into play: orbital angle, bar position angle, primary galaxy mass...



Conclusion

- Fraction of barred galaxies in Illustris is in the lower values of observations, increases with galaxy mass
- Tidally induced bars: bar is created right after the pericenter passage of an encounter
- Prograde orbits is the preferred way to create a bar or increase the bar strength
- The stronger the interaction, the stronger the created bar strength, but many other parameters can come into play

Thank you!